

Stable isotope geochemistry



Franck Poitrasson
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Suggested readings:

- Allègre, C.J., 2008. Isotope geology. Cambridge University Press, Cambridge, 512 pp.
- Hoefs, J., 2004. Stable isotope geochemistry. Springer-Verlag, Berlin, 244 pp.
- Criss, R.E., 1999. Principles of stable isotope distribution. Oxford University Press, Oxford, 254 pp.
- Reviews in Mineralogy (and Geochemistry) on Stable Isotopes (vol. 16, 1986 & vol. 43, 2001)

Isotopes

- Atoms with the same number of protons, but a different number of neutrons:

→ relate to composition of the nucleus of atoms



Carbon-12
6 protons
6 neutrons
6 electrons



Carbon-13
6 protons
7 neutrons
6 electrons



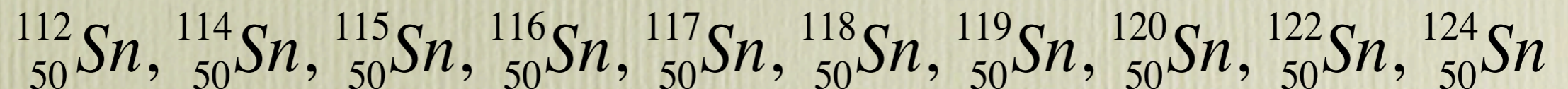
Carbon-14
6 protons
8 neutrons
6 electrons

- The number of stable isotopes may vary depending on the chemical element:
- From one: ${}_{11}^{23}\text{Na}$, ${}_{41}^{93}\text{Nb}$, ${}_{67}^{165}\text{Ho}$, ${}_{79}^{197}\text{Au}$ (21 naturally occurring)
- To several: ${}_{29}^{63}\text{Cu}$, ${}_{29}^{65}\text{Cu}$

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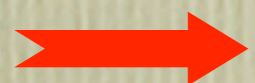
- To several: ${}_{29}^{63}\text{Cu}$, ${}_{29}^{65}\text{Cu}$



 Isotope geochemistry requires at least two isotopes!

Periodic chart of elements

H																	He																												
Li	Be											B	C	N	O	F	Ne																												
Na	Mg											Al	Si	P	S	Cl	Ar																												
K	Ca	Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As	Se	Br	Kr																												
Rb	Sr	Y	Zr	Nb	Mo	Tc	Ru	Rh	Pd	Ag	Cd	In	Sn	Sb	Te	I	Xe																												
Cs	Ba	La	Hf	Ta	W	Re	Os	Ir	Pt	Au	Hg	Tl	Pb	Bi	Po	At	Rn																												
Fr	Ra	Ac	Rf	Ha																																									
<table border="1"> <tr> <td>Ce</td> <td>Pr</td> <td>Nd</td> <td>Pm</td> <td>Sm</td> <td>Eu</td> <td>Gd</td> <td>Tb</td> <td>Dy</td> <td>Ho</td> <td>Er</td> <td>Tm</td> <td>Yb</td> <td>Lu</td> </tr> <tr> <td>Th</td> <td>Pa</td> <td>U</td> <td>Np</td> <td>Pu</td> <td>Am</td> <td>Cm</td> <td>Bk</td> <td>Cf</td> <td>Es</td> <td>Fm</td> <td>Md</td> <td>No</td> <td>Lr</td> </tr> </table>																		Ce	Pr	Nd	Pm	Sm	Eu	Gd	Tb	Dy	Ho	Er	Tm	Yb	Lu	Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm	Md	No	Lr
Ce	Pr	Nd	Pm	Sm	Eu	Gd	Tb	Dy	Ho	Er	Tm	Yb	Lu																																
Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm	Md	No	Lr																																
 Only one stable isotope		 Naturally occurring, may be analyzed by MC-ICP-MS																																											
 Difficult to analyze by MC-ICP-MS																																													



Over 60 elements suitable for stable isotope studies

The “XVIII^e century view”:

- Atoms with different number of protons have different names because they have different chemical and physical properties

The “XVIII^e century view”:

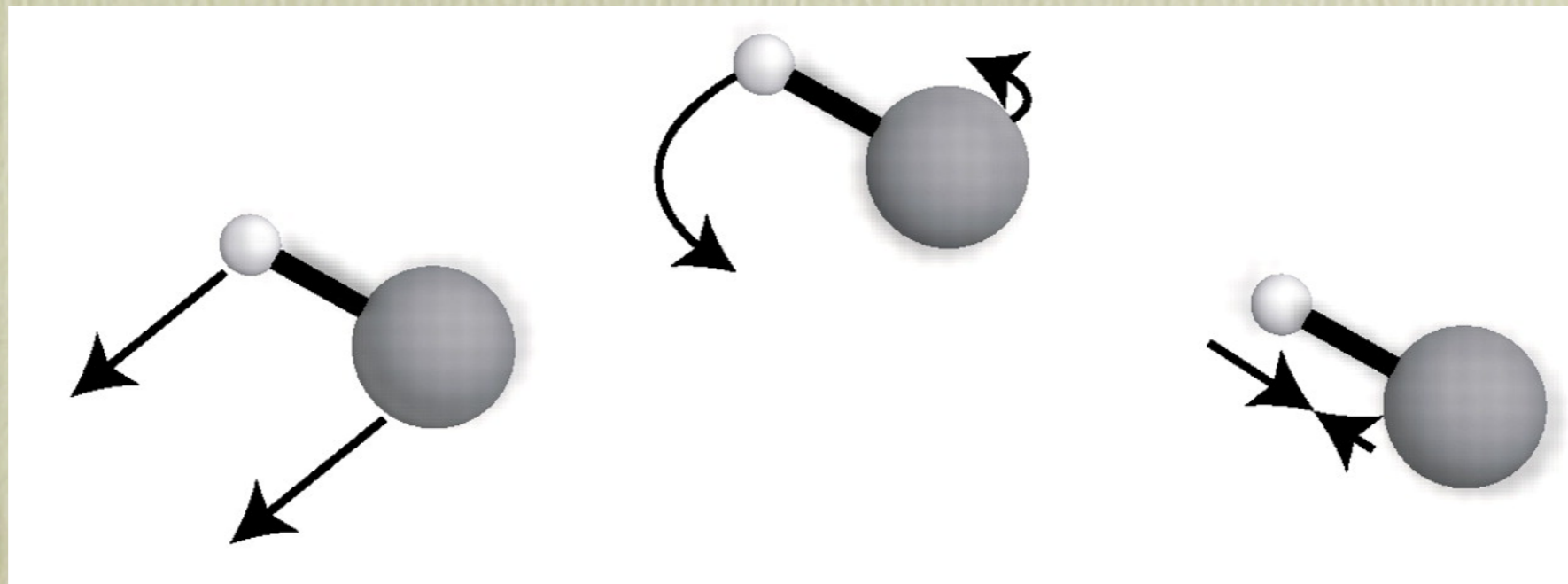
- Atoms with different number of protons have different names because they have different chemical and physical properties

The XX^e century view:

- In detail, atoms with the same number of protons but different number of neutrons (isotopes) have slightly different chemical and physical properties due to slight mass difference of their host molecules...

How is it possible?

- Gaseous, aqueous molecules and mineral lattices are vibrating continuously, and the vibration frequencies will depend on the isotopic composition (weight) of the molecule
- They are three vibrating modes:



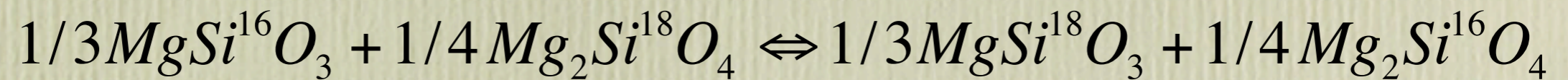
Schauble.,
2004, RiMG

translational

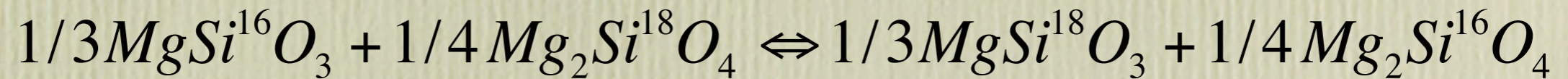
rotational

vibrational

- The effect of these different levels of energy on isotopic signatures may be observed when exchanges between two reservoirs occur, i.e., through an exchange reaction.



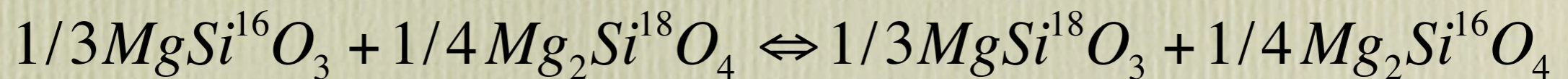
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The equilibrium constant is (assuming ideal mixing):

$$K = \frac{\left({}^{18}O_{opx}^3 \right)^{1/3} \cdot \left({}^{16}O_{ol}^4 \right)^{1/4}}{\left({}^{16}O_{opx}^3 \right)^{1/3} \cdot \left({}^{18}O_{ol}^4 \right)^{1/4}} = \frac{\left[\left({}^{18}O/{}^{16}O \right)_{opx}^3 \right]^{1/3}}{\left[\left({}^{18}O/{}^{16}O \right)_{ol}^4 \right]^{1/4}} = \frac{\left({}^{18}O/{}^{16}O \right)_{opx}}{\left({}^{18}O/{}^{16}O \right)_{ol}} = \alpha_{opx-ol}$$

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Each molar quantity above may be calculated through partition functions that describes all possible energy states of the substance:

$$Q = Q_{tr} \cdot Q_{rot} \cdot Q_{vib}$$

Notations and conventions:

- So far, we only used the isotopic fractionation factor α between two substances A and B, defined as:

$$\alpha_{A-B} = \frac{\left(N^*/N\right)_A}{\left(N^*/N\right)_B} \quad \text{where } N^* \text{ and } N \text{ are the number of atoms of the heavy and light isotopes, respectively}$$

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$$\delta_A = 1000 \cdot \left[\frac{\left(N^*/N\right)_A}{\left(N^*/N\right)_{std}} - 1 \right] \quad \text{expressed in } \text{‰}$$

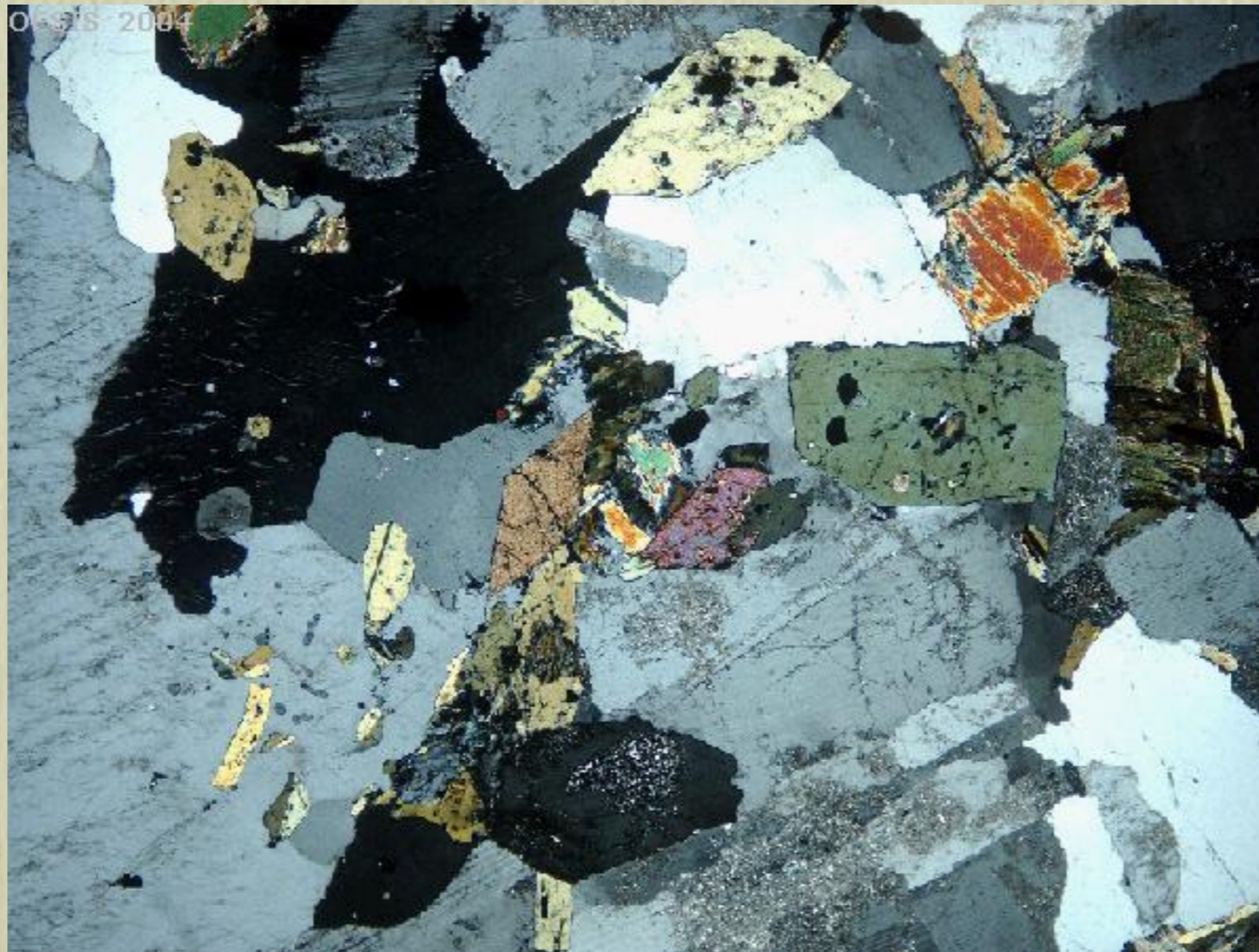
- It may be shown that α and δ are linked:

$$\alpha_{A-B} = \frac{\left(N^*/N\right)_A}{\left(N^*/N\right)_B} = \frac{1000 + \delta_A}{1000 + \delta_B}$$

But how do we measure these
stable isotope ratio?

Sample preparation: some basic principles

- For bulk solid samples, the starting material should be large enough to average local heterogeneity



Granodiorite: cm-scale minerals are common

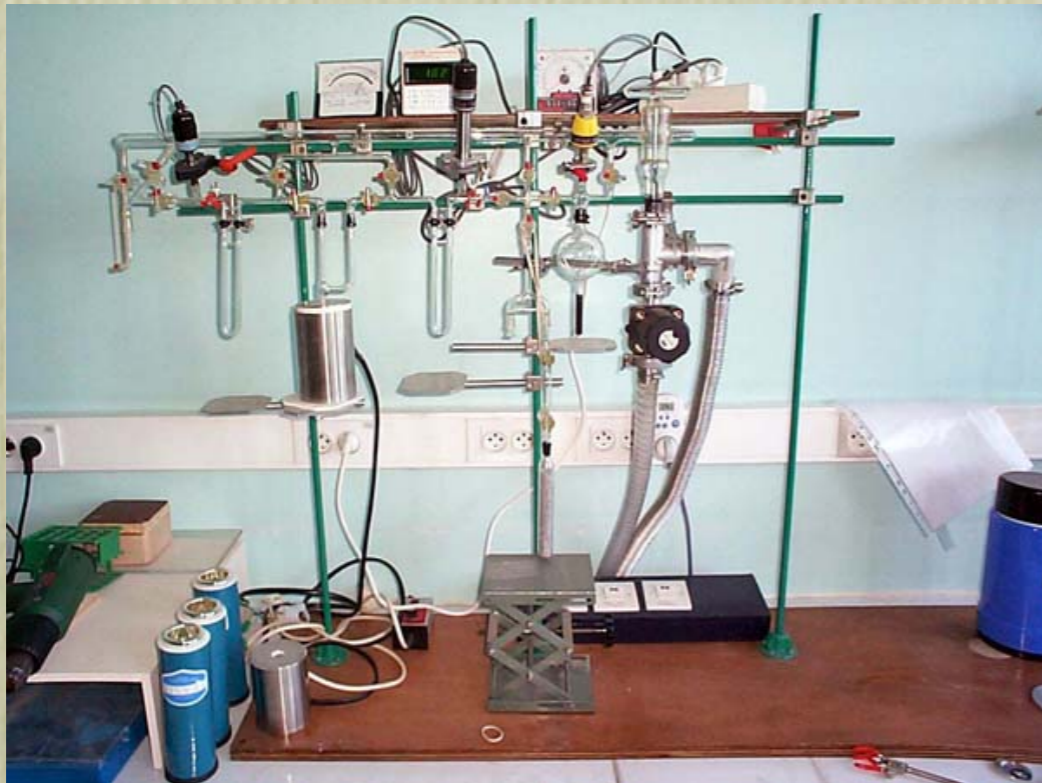
Sample preparation: some basic principles

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- Isotope ratios are frequently measured on gases

Element	Gas
H	H ₂
C	CO ₂ , CO
N	N ₂
O	CO ₂ , CO, O ₂
S	SO ₂ , SF ₆
Si	SiF ₄

Sample preparation: some basic principles

- For bulk solid samples, the starting material should be large enough to average local heterogeneity
- Isotope ratios are frequently measured on gases
- The element of interest should be purified, i.e., separated from its matrix



H extraction line



O extraction line from silicates

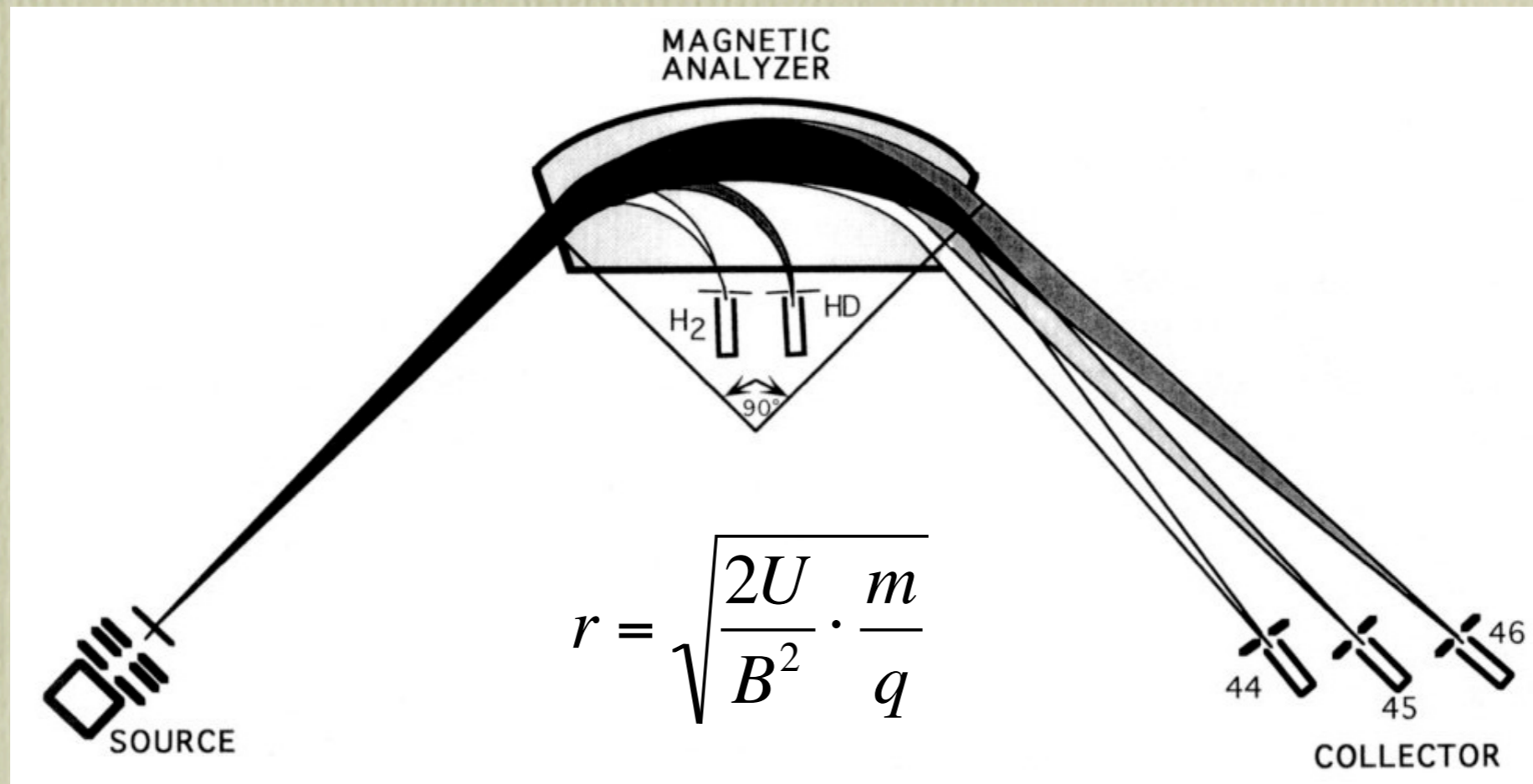
Sample preparation: some basic principles

- For bulk solid samples, the starting material should be large enough to average local heterogeneity
- Isotope ratios are frequently measured on gases
- The element of interest should be purified, i.e., separated from its matrix
- Methods should be clean to avoid contamination. The mg level is the typical sample size.
- 100% yield during purification is required

Gas source mass spectrometer



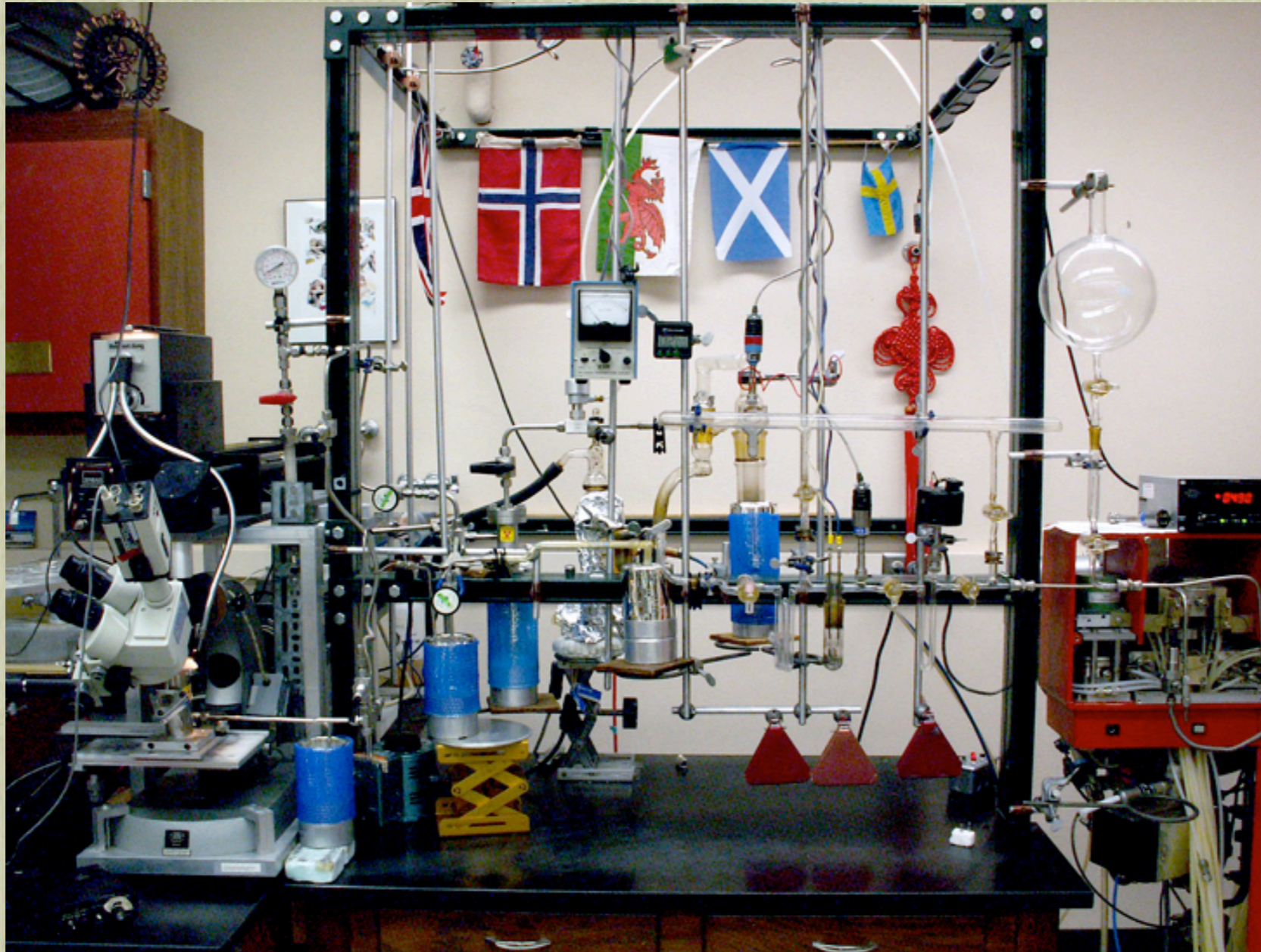
Gas source mass spectrometer



where: U is the accelerating voltage, B the magnetic field, m/q the mass/charge ratio of the ion and r the radius of its circular path in the magnetic field

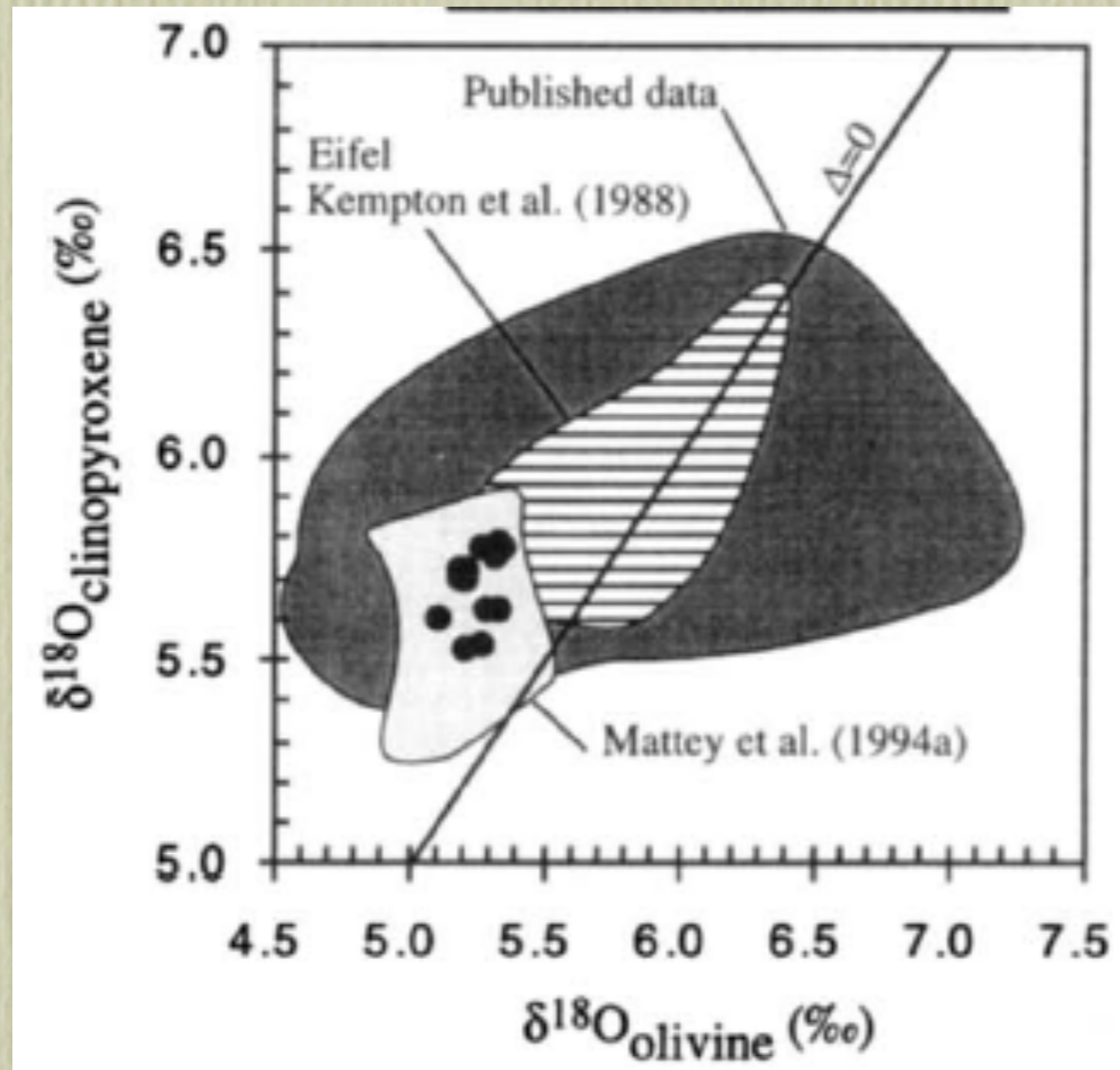
➡ The lighter ions will turn faster!

Laser fluorination for mineral analysis



University of Madison-Wisconsin laboratory photo

Benefit of laser fluorination for mineral analysis:



Mattey et al., 1994: Laser fluorination data
Others: conventional data

Chazot et al., 1997, GCA

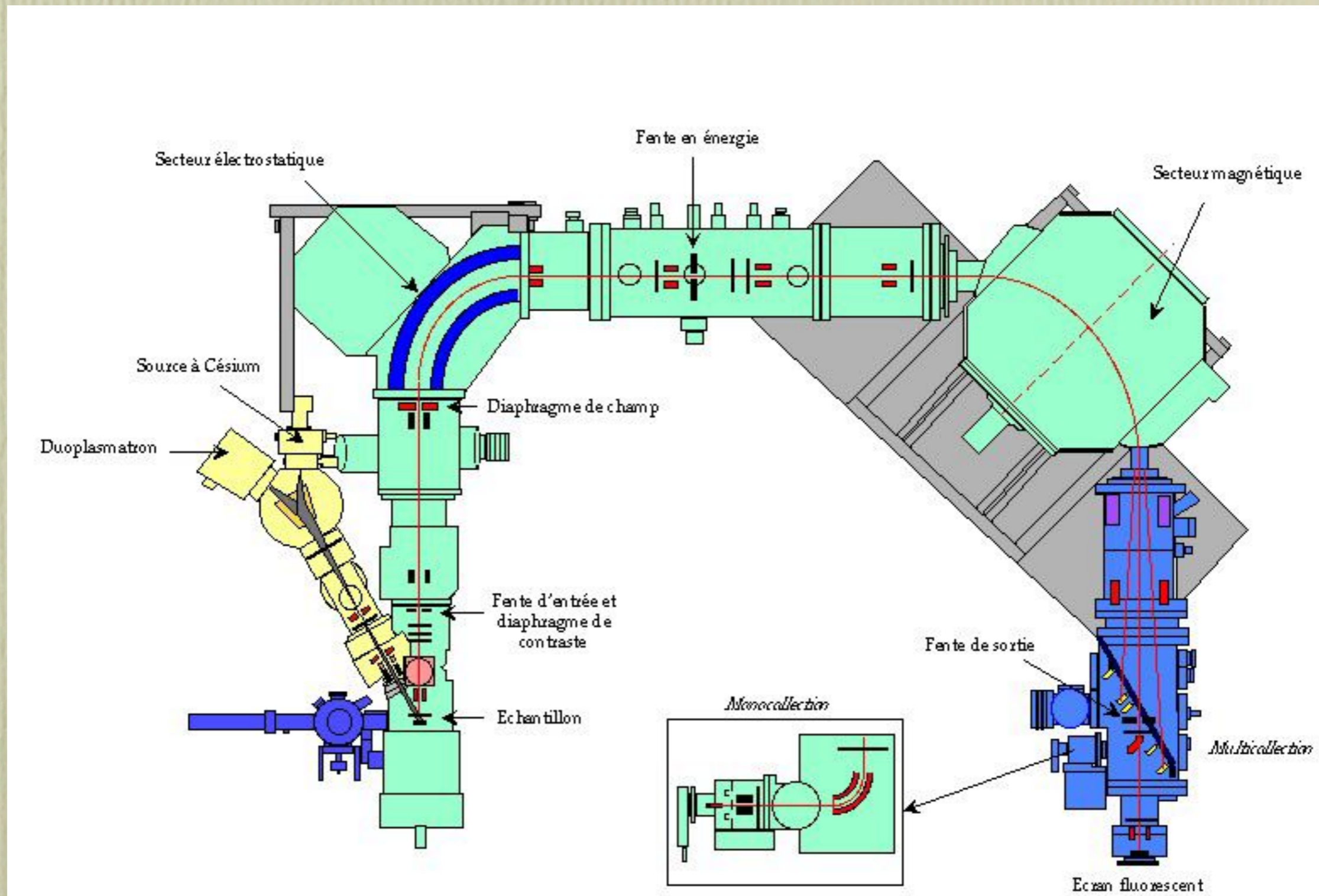
➔ The range in O isotope composition previously reported was an analytical artifact!

In situ light (H, C, O...) stable isotope measurements: ion microprobe



CNRS-CRPG Nancy Laboratory photo

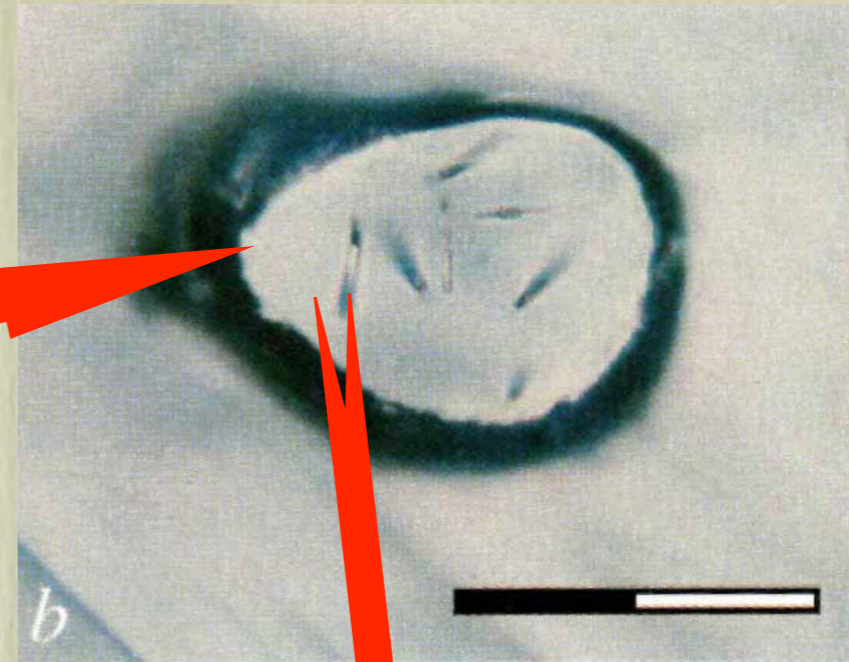
In situ light (H, C, O...) stable isotope measurements: ion microprobe



Ion microprobe: very useful for very small samples



3,8 Ga Akilia gneiss, Greenland



Apatite
(scale: 20 μm)

Carbon-rich inclusion with
a life-like $\delta^{13}\text{C}$ signature

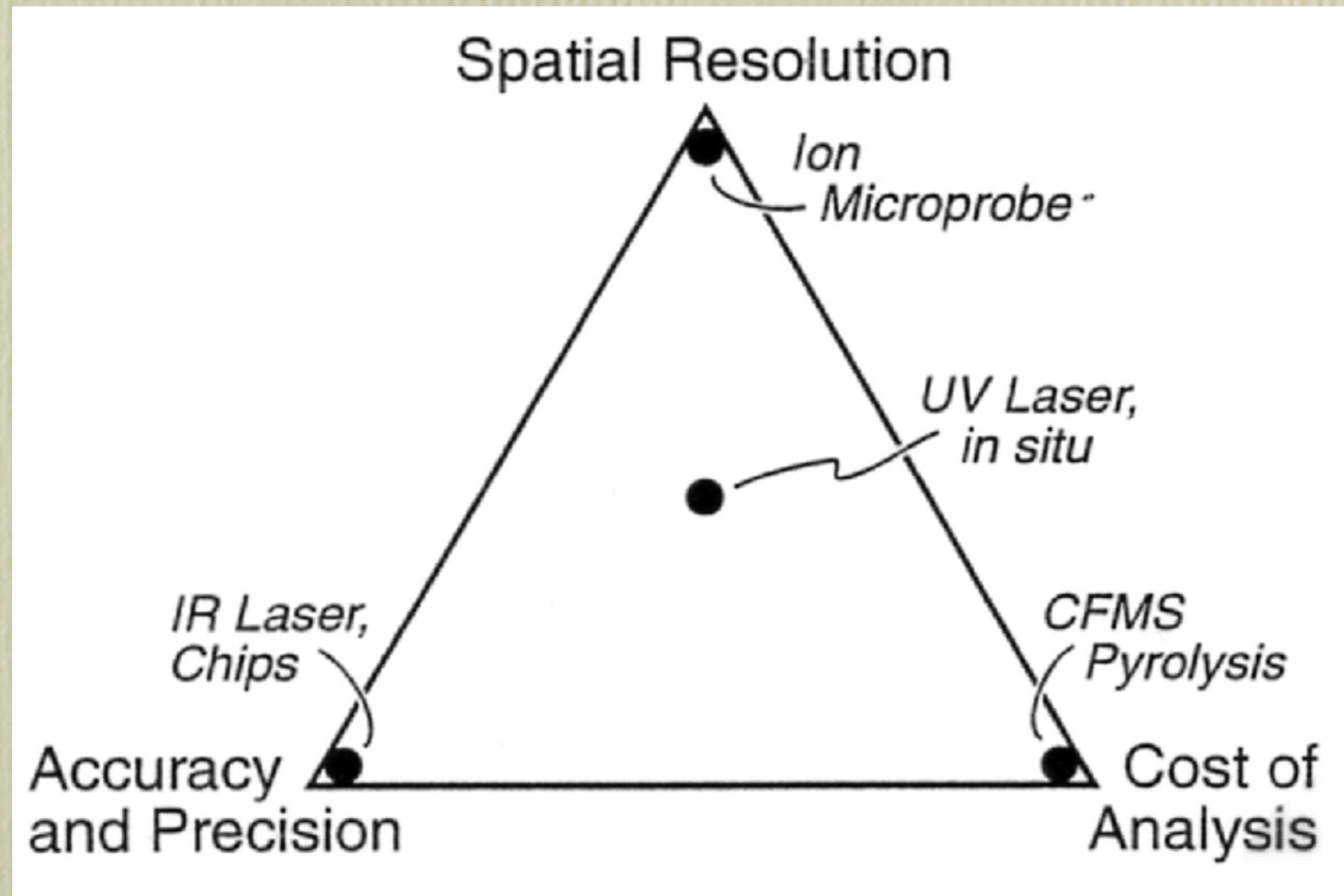


After Mosjzsis et al., 1996, Nature

Characteristics of various techniques for light isotope analysis (Valley et al., 1998, RiMG)

Technique	Isotopes, Minerals	Sample Prep.	Typical Sample Size	Spatial Resolution		Precision (1 sd)	References
CONVENTIONAL TECHNIQUES							
Ni reaction vessels	O silicates, oxides	Powder or chips	10-20 mg	-	-	0.2‰ ^A	1
Phosphoric acid	O,C carbonates	Powder or chips	> 10µg	-	-	0.03-0.1‰ ^B	2
Combustion	C graphite, diamonds	Powder or chips	> 1µg	-	-	0.03‰	3
Combustion	S	Powder or chips	> 1 mg	-	-	0.1‰	4
Fusion	H solids	Powder or chips	50-100 mg	-	-	1-2‰	5
LASER PROBE / MASS-SPECTROMETER							
Laser probe, IR	O silicates, oxides	Powder or chips	≥ 0.5 mg	500	µm	0.07‰	6,7
Laser probe, IR	O silicates, oxides	In situ	≥ 0.5 mg	500	µm (5mm) ^C	0.3-0.5‰	6
Laser probe, IR	S	In situ, chips	≥0.1 mg	200	µm	0.2‰	8
Laser probe, IR	O,C carbonates	In situ	≥0.5 mg	500	µm	^D	9
Laser probe, IR	O phosphates	Powder or chips	≥0.5 mg	500	µm	0.1‰	10
Laser probe, UV	O silicates, oxides	In situ	~ 0.5 mg	~500	µm	0.1‰	11
Laser probe, UV	O phosphates	In situ		300x400	µm	0.4‰	12
ION MICROPROBE / SIMS							
Ion probe	O oxides (conductors)	In situ	0.4 ng	8	µm	0.5-1‰	13
Ion probe	O silicates or conductors	In situ	5 ng	20	µm	0.5-1‰	14
Ion probe	O carbonates	In situ	5 ng	20	µm	0.5-1‰	15
Ion probe	S	In situ	1-5 ng	10-20	µm	0.25-1‰	16
Ion probe	H	In situ	5-10 ng	20-30	µm	10‰	17
Ion probe	C	In situ	1-5 ng	10-30	µm	0.5-1‰	18

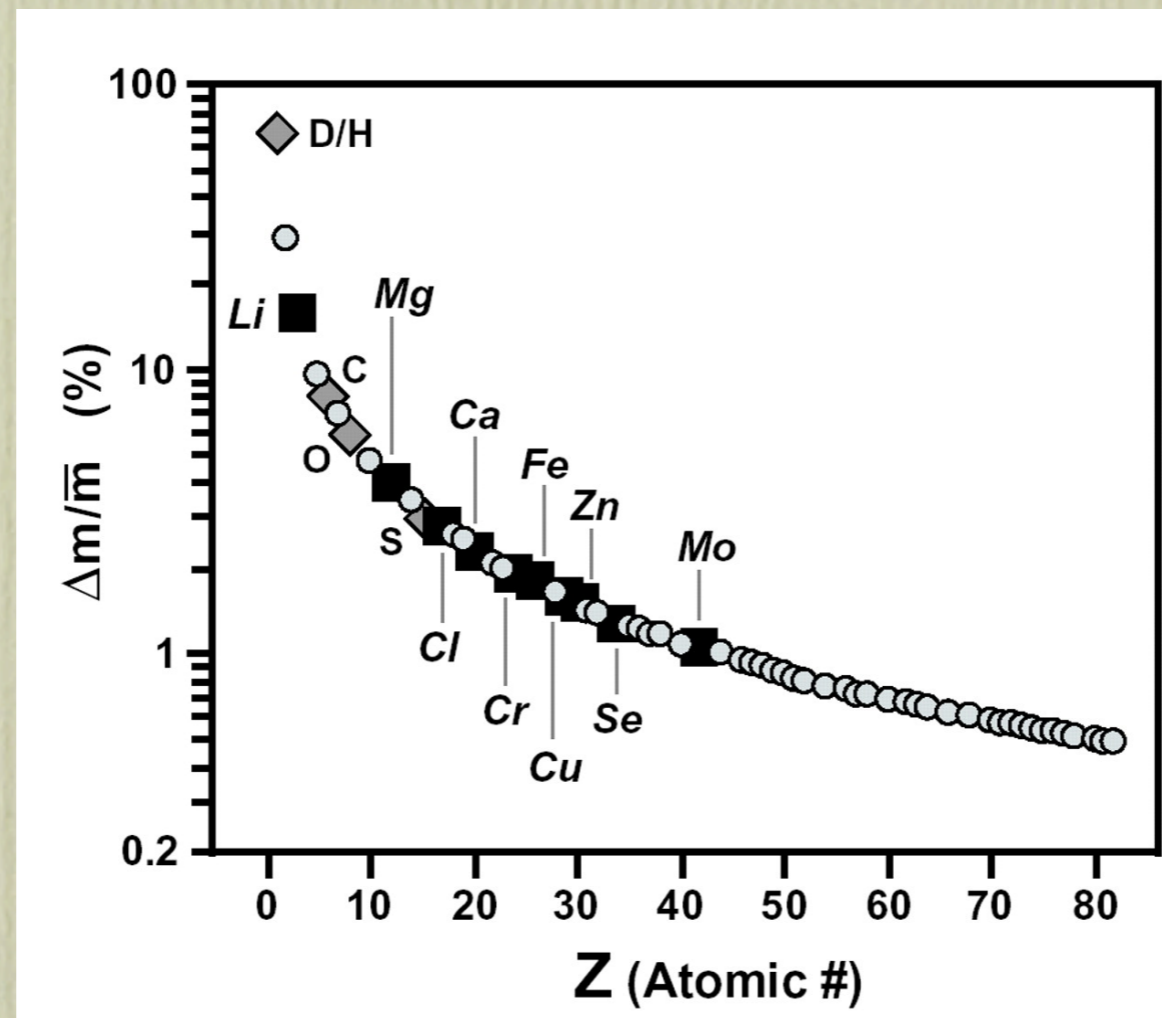
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Factors affecting equilibrium stable isotope fractionation:

Factors affecting equilibrium stable isotope fractionation:

- Mass difference between isotopes



In general, the larger the relative mass difference between isotopes of a given element, the larger the stable isotope fractionation in nature ... but they are many exceptions...

Factors affecting equilibrium stable isotope fractionation:

- Mass difference between isotopes
- Nature of the interatomic bonds

C-C bonds in diamond are more covalent than in graphite, and therefore of higher energy. Thus:

$$\alpha_{dia-gra} = \frac{\left({}^{13}\text{C}/{}^{12}\text{C} \right)_{dia}}{\left({}^{13}\text{C}/{}^{12}\text{C} \right)_{gra}} > 1$$

Factors affecting equilibrium stable isotope fractionation:

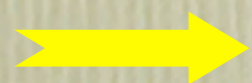
- Mass difference between isotopes
- Nature of the interatomic bonds
- Coordination chemistry



**Triangular
Higher energy**

**tetrahedral coordination
Lower energy**

$$\alpha_{B(OH)_4^- - B(OH)_3} = \frac{\left(\frac{{}^{11}B}{{}^{10}B} \right)_{B(OH)_4^-}}{\left(\frac{{}^{11}B}{{}^{10}B} \right)_{B(OH)_3}} < 1$$



A possible approach to measure ocean's paleo-pH?

Factors affecting equilibrium stable isotope fractionation:

- Mass difference between isotopes
- Nature of the interatomic bonds
- Coordination chemistry
- Redox state of elements

S^{2-} in sulfide species, S^0 in native sulfur, S^{4+} in SO_2 and S^{6+} in sulfate

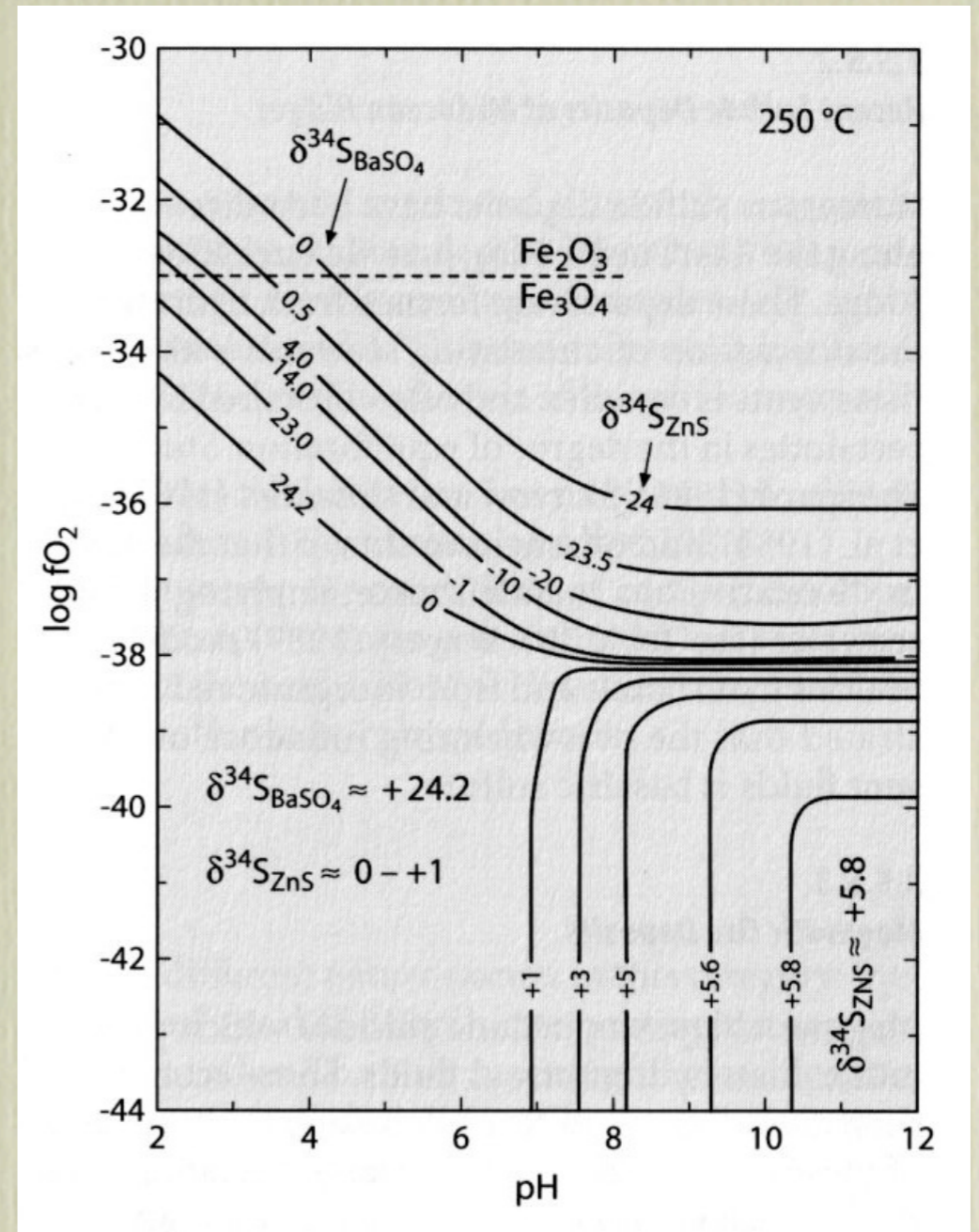
and:

$$\alpha_{sulfates-sulfides} = \frac{\left(\frac{{}^{34}S}{{}^{32}S} \right)_{sulfates}}{\left(\frac{{}^{34}S}{{}^{32}S} \right)_{sulfides}} > \alpha_{SO_2-sulfides} = \frac{\left(\frac{{}^{34}S}{{}^{32}S} \right)_{SO_2}}{\left(\frac{{}^{34}S}{{}^{32}S} \right)_{sulfides}} > 1$$

Application to estimate the formation conditions of ore deposits



Madan, Bulgaria

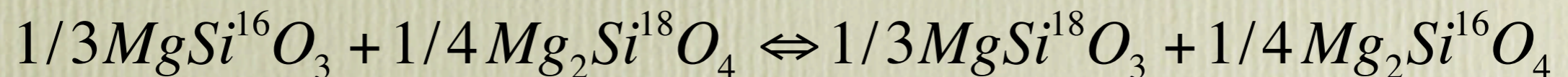


After Ohmoto,
1972, Econ. Geol.

Factors affecting equilibrium stable isotope fractionation:

- Mass difference between isotopes
- Nature of the interatomic bonds
- Coordination chemistry
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- Temperature

We have seen that for a reaction like:

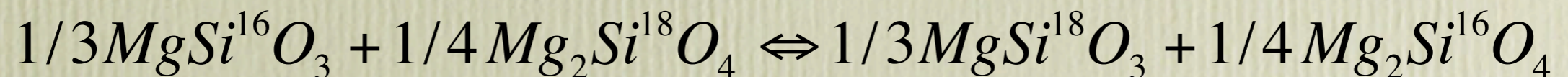


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Factors affecting equilibrium stable isotope fractionation:

- Mass difference between isotopes
- Nature of the interatomic bonds
- Coordination chemistry
- Redox state of elements
- Temperature

We have seen that for a reaction like:



$\alpha = K$, if the stoichiometric coefficients of the reaction are well chosen. A more generic relationship is: $\alpha = K^{1/n}$

where n is the number of atoms exchanged.

The equilibrium constant K is linked to T by:

$$\Delta G = -RT \ln K$$

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For isotope exchanges, temperature dependence is more precisely described using statistical mechanics that allows to compute partition functions (Q):

$$(Q^*/Q) = \left(\frac{m^*}{m}\right)^{3n/2} \frac{\sigma}{\sigma^*} \prod_i \frac{\nu_i^* e^{-U_i^*/2}}{\nu_i e^{-U_i/2}} \frac{1 - e^{-U_i}}{1 - e^{-U_i^*}}$$

where * stands for the heavy isotope, m for mass, σ for symmetry number, ν for vibration frequency and:

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$$U_i = h\nu_i / kT$$

where h is the Planck's constant, k is the Boltzmann's constant, i is a running index of vibrational modes and T is the temperature in $^{\circ}\text{K}$.

Use of these equations and experimental work show that:

$$1000 \ln \alpha = A/T^2 + B$$

at $T \geq 500-600^\circ\text{C}$ for minerals and $\geq 100^\circ\text{C}$ for gases, and

$$1000 \ln \alpha = A/T + B$$

at lower temperatures

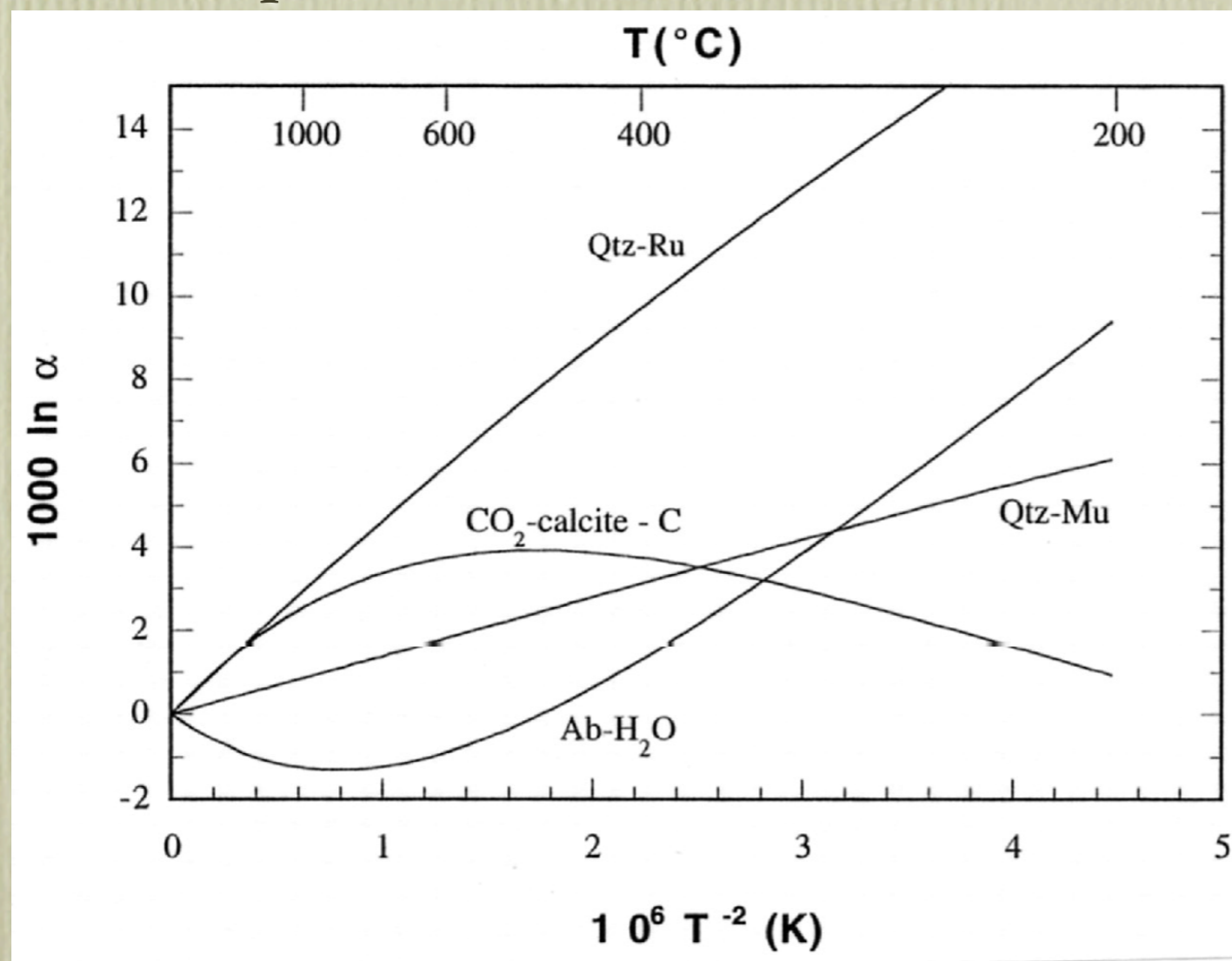
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at lower temperatures, but other relationships and exceptions exist:



Oxygen and carbon (- C) fractionation factors as a function of T

Chacko et al.,
2001, RiMG

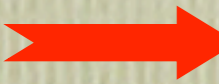
➔ Thermometry possible in simple and well constrained cases

Kinetic stable isotope fractionation

- Occurs when not enough time is given for exchange reactions to reach completion (e.g., at low T , in biological processes, if diffusion in solids is required), or
- When products and reactants are rapidly separated (e.g., through a phase change interface, through a biological cell)

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- When products and reactants are rapidly separated (e.g., through a phase change interface, through a biological cell)

 In those situations, lightest isotopes are enriched in the reaction products because of their fastest velocity and their lower activation energy required for reactions to proceed

Kinetic stable isotope fractionation for diffusing gases:

For ideal gases: $E_c = E_c^*$ and $E_c = 1/2mV^2 (= 3/2kT)$

where * stands for the heavy gaseous species. As a result:

$$\frac{V}{V^*} = \sqrt{\frac{m^*}{m}} = \alpha_1 > 1 \quad \text{where } \alpha \text{ is the isotopic fractionation factor}$$

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$$\frac{D}{D^*} = \sqrt{\frac{1/m + 1/m_{dm}}{1/m^* + 1/m_{dm}}} = \alpha_2$$

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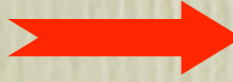
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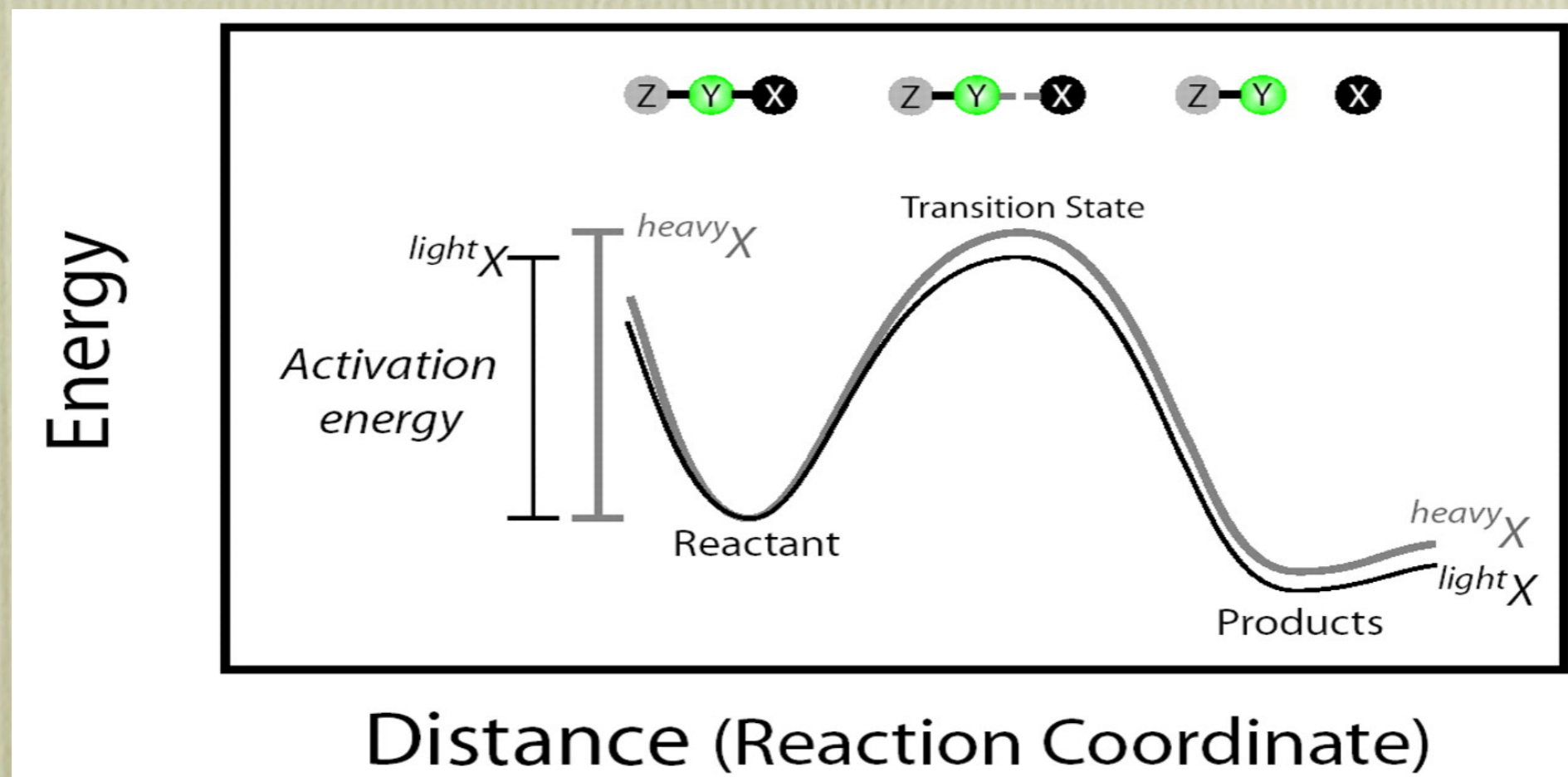
We will have: $\alpha_1 > \alpha_2 > 1$  Medium-limited diffusion will minimize kinetic isotope fractionation factors

Kinetic stable isotope fractionation during a chemical reaction:

In simple cases where only one molecule dissociates, a simplified approach to calculate isotopic effects on reaction rates is used, in which only the energy barrier between the reactants and a single rate-limiting transition-state is considered:

$$p(E_A) = \exp(-E_A / kT)$$

where $p(E_A)$ is the probability of a molecule having enough energy to reach the transition-state of activation energy E_A .



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- The isotopic fractionation factor may be expressed as the difference between reduced partition functions of substance A and B:

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Notations and conventions:

- An other useful approximation, possible for most stable isotope fractionation observed in nature:

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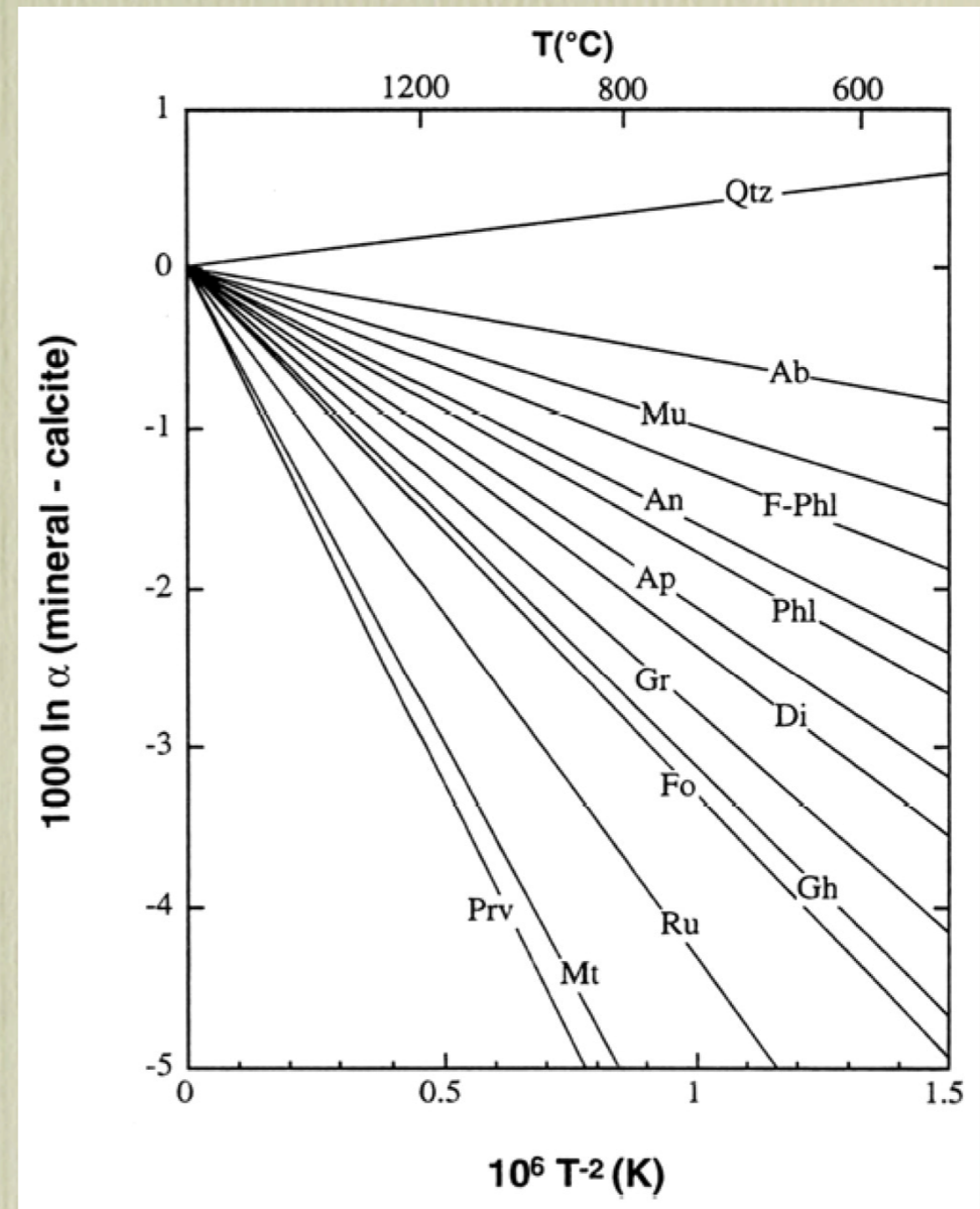
- These reduced partition function computed from statistical mechanics methods are commonly in the form:

$$1000 \cdot \ln \beta_A = a \cdot x + b \cdot x^2 + c \cdot x^3$$

where $x = 10^6 / T^2$ with T in K and a, b and c are tabulated factors

Example of reduced partition function for oxygen isotopes in minerals

Mineral	$1000 \ln \beta$
Calcite	$11.781 x - 0.420 x^2 + 0.0158 x^3$
Quartz	$12.116 x - 0.370 x^2 + 0.0123 x^3$
Albite	$11.134 x - 0.326 x^2 + 0.0104 x^3$
Muscovite	$10.766 x - 0.412 x^2 + 0.0209 x^3$
Anorthite	$9.993 x - 0.271 x^2 + 0.0082 x^3$
Phlogopite	$9.969 x - 0.382 x^2 + 0.0194 x^3$
F-phlogopite	$10.475 x - 0.401 x^2 + 0.0203 x^3$
Diopside	$9.237 x - 0.199 x^2 + 0.0053 x^3$
Forsterite	$8.236 x - 0.142 x^2 + 0.0032 x^3$
Rutile	$7.258 x - 0.125 x^2 + 0.0033 x^3$
Magnetite	$5.674 x - 0.038 x^2 + 0.0003 x^3$



Chacko et al.,
2001, RiMG

Notations and standards used as reference for delta values

$$\delta_A = 1000 \cdot \left[\frac{\left(N^* / N \right)_A}{\left(N^* / N \right)_{std}} - 1 \right] \quad \text{in } \text{‰}$$

Standard	Ratio	Accepted value ($\times 10^6$) (with 95% confidence interval)	Source
SMOW	D/H	155.76 \pm 0.10	Hagemann et al. (1970)
	$^{18}\text{O}/^{16}\text{O}$	2,005.20 \pm 0.43	Baertschi (1976)
	$^{17}\text{O}/^{16}\text{O}$	373 \pm 15	Nier (1950), corrected by Hayes(1983)
PDB	$^{13}\text{C}/^{12}\text{C}$	-11,237.2 \pm 2.9	Craig (1957)
	$^{18}\text{O}/^{16}\text{O}$	2,067.1 \pm 2.1	
	$^{17}\text{O}/^{16}\text{O}$	379 \pm 15	
Air nitrogen	$^{15}\text{N}/^{14}\text{N}$	3,676.5 \pm 8.1	Junk and Svec (1958)
Canyon Diablo Troilite (CDT)	$^{34}\text{S}/^{32}\text{S}$	45,004.5 \pm 9.3	Jensen and Nakai (1962)

Mixing relationships

we take: $R = N^* / N$

- Closed system (e.g., minerals in a rock):

$$R_M = X_A \cdot R_A + (1 - X_A) \cdot R_B \quad \text{in delta:} \quad \delta_M = X_A \cdot \delta_A + (1 - X_A) \cdot \delta_B$$

where δ_M stands for the isotopic composition of a mixture of components A and B and X_A is the molar fraction of component A.

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$$R = \alpha \cdot R_0 \quad \text{and since:} \quad \frac{R}{R_0} = \frac{1 + \frac{\delta}{1000}}{1 + \frac{\delta_0}{1000}} \approx 1 + \frac{(\delta - \delta_0)}{1000}$$

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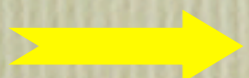
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 In the “open system” situation, there is no need to consider mass balance constraints, only the isotopic fractionation factor and the isotopic composition of the big reservoir will determine the isotopic composition of the small reservoir.

A useful specific open-system case: the Rayleigh isotopic fractionation

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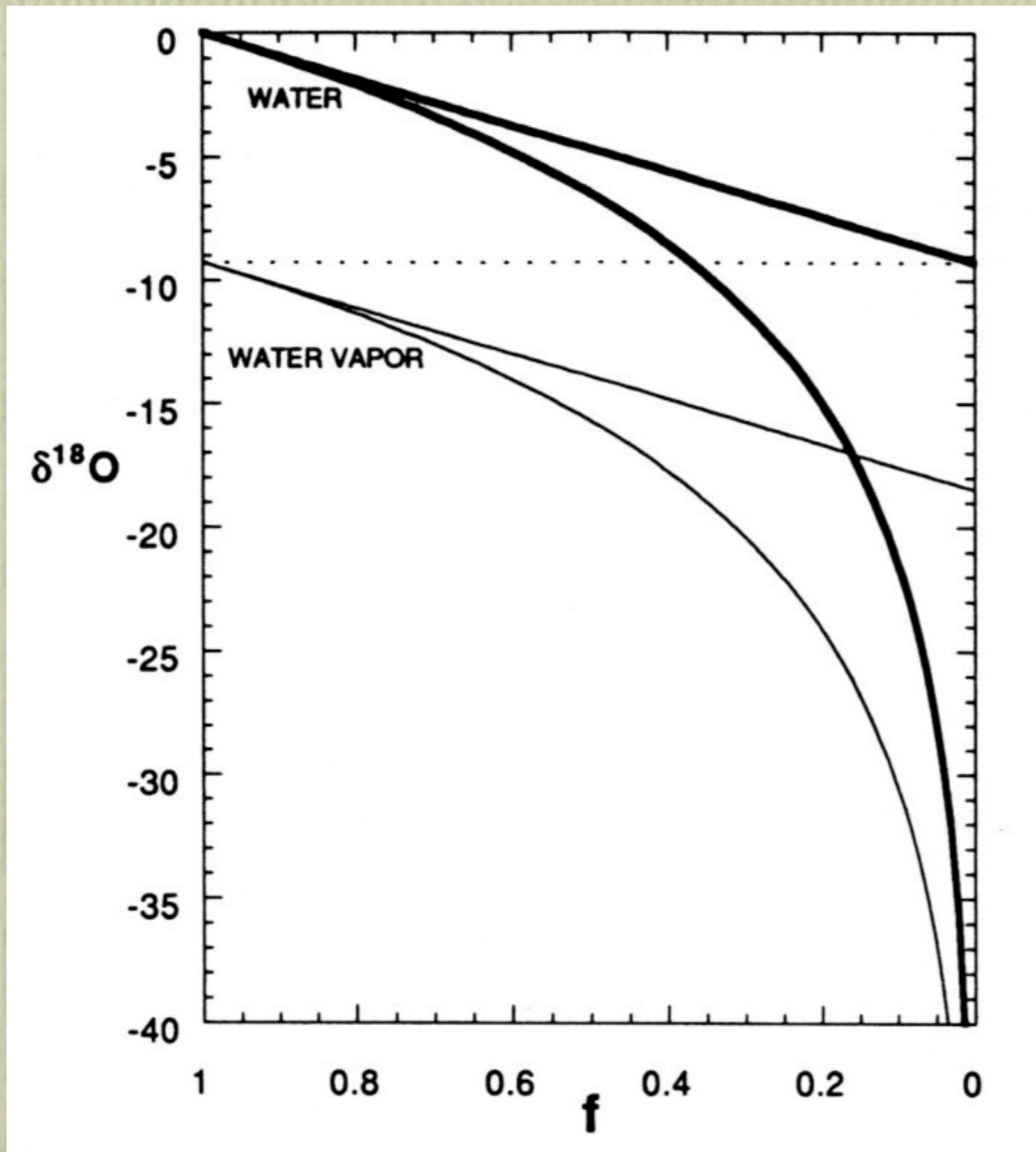
$$R = R_i \cdot f^{(\alpha-1)}$$

where R_i is the isotopic ratio of the reservoir at the beginning of the process, when the fraction remaining, $f = 1$.
 α is the fractionation factor.

Using the delta notation, this expression may be rearranged as:

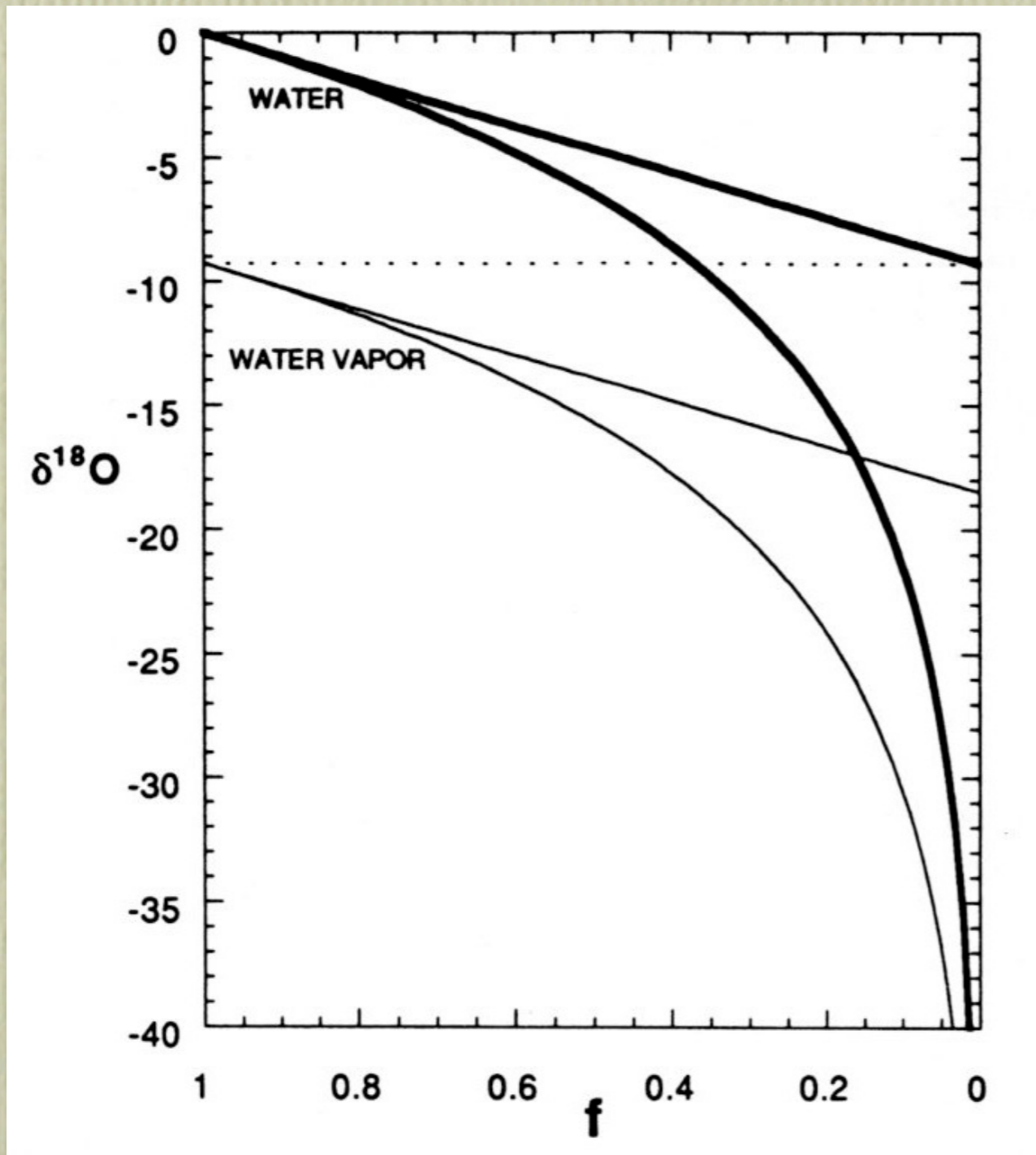
$$\delta = \delta_i + 1000 \cdot (\alpha - 1) \cdot \ln f$$

Example of water condensation effect on oxygen isotopes:



- Closed system (straight lines): the last condensate reach the isotopic composition of the starting vapor

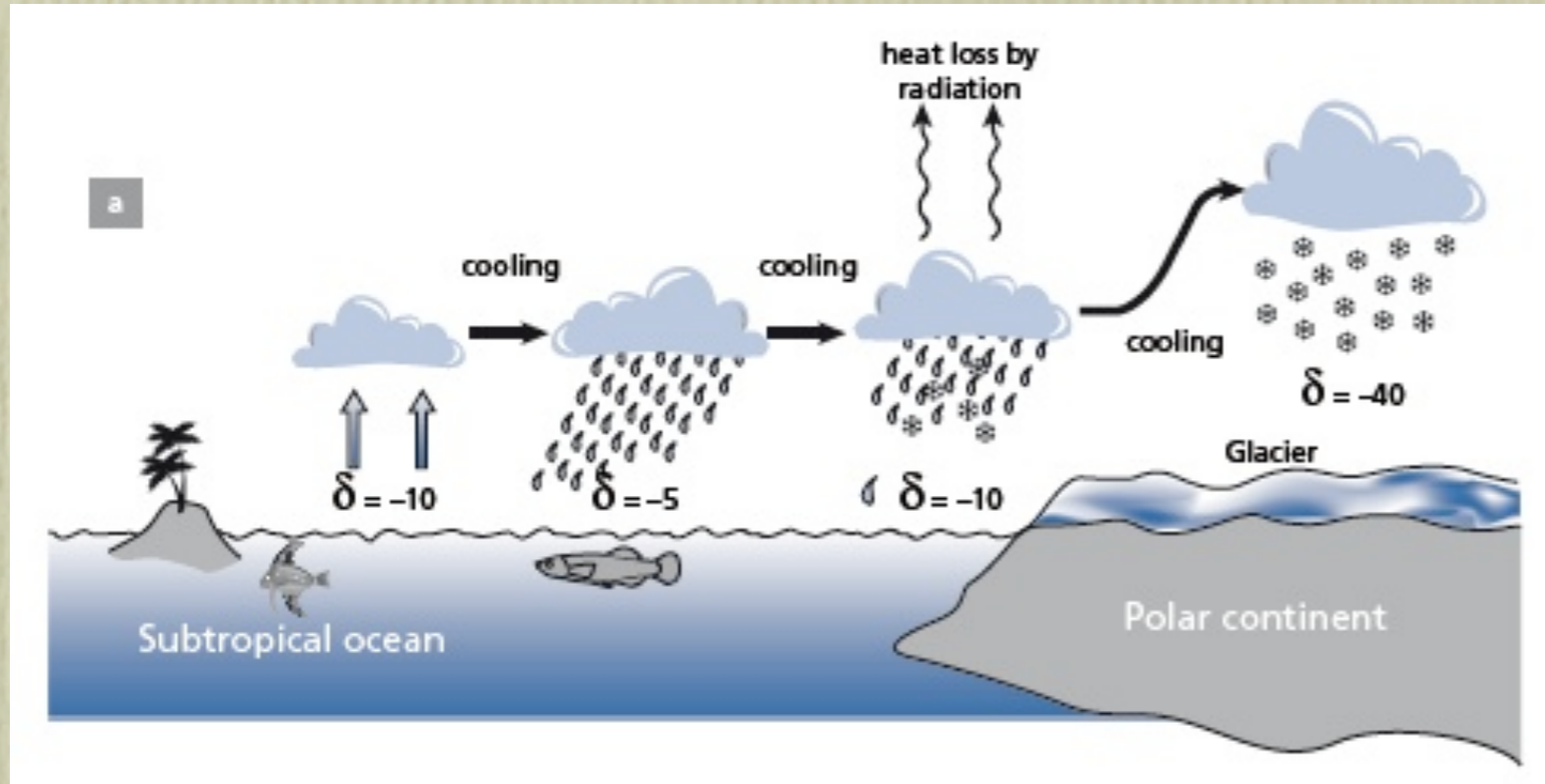
Example of water condensation effect on oxygen isotopes:



- Closed system (straight lines): the last condensate reach the isotopic composition of the starting vapor
- Open system (Rayleigh distillation, curves): the last condensate reach much lighter oxygen isotope composition

Note: α is constant in both cases

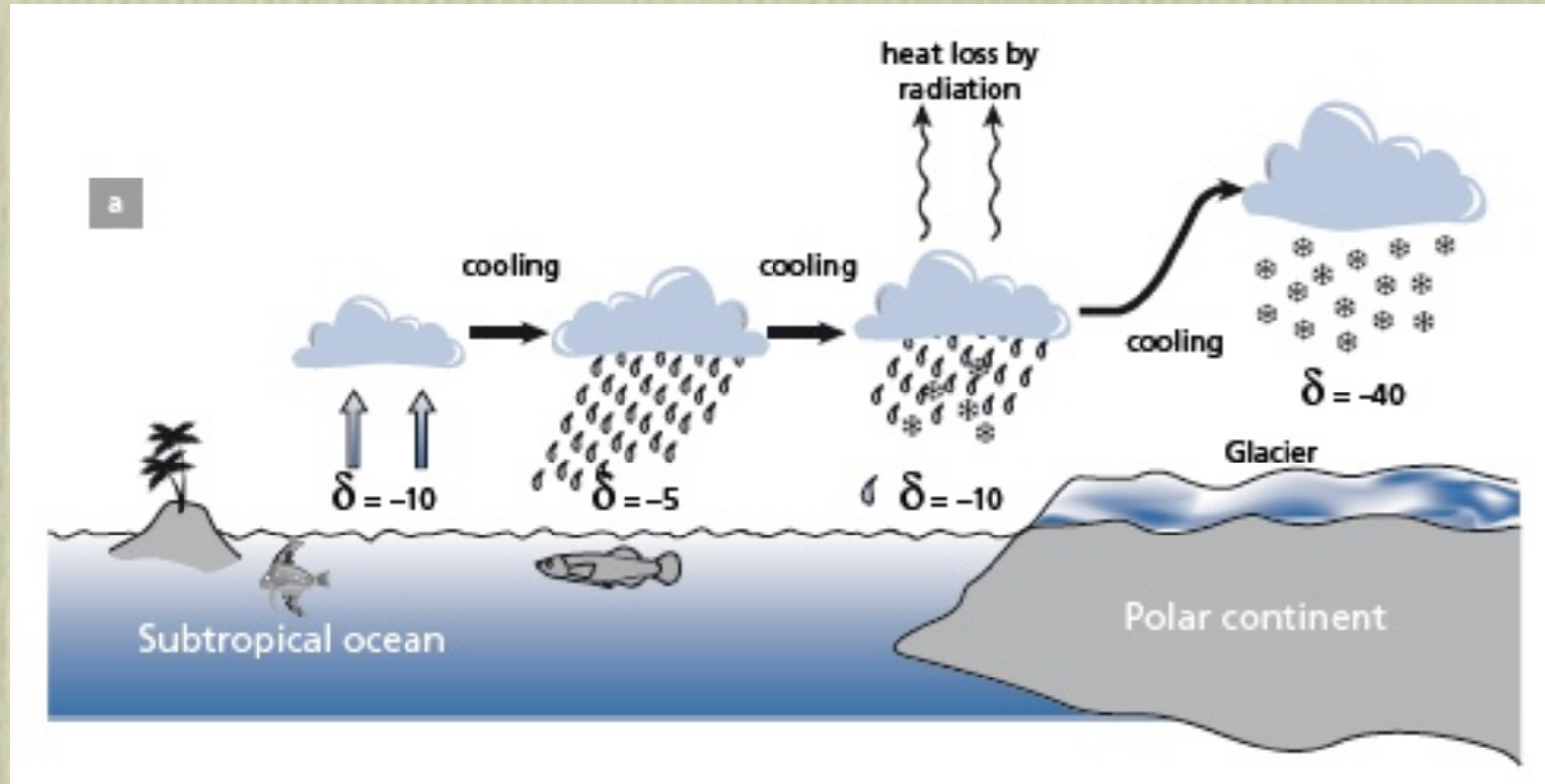
Application to rainwater $\delta^{18}\text{O}$ (‰) from the tropics to the poles



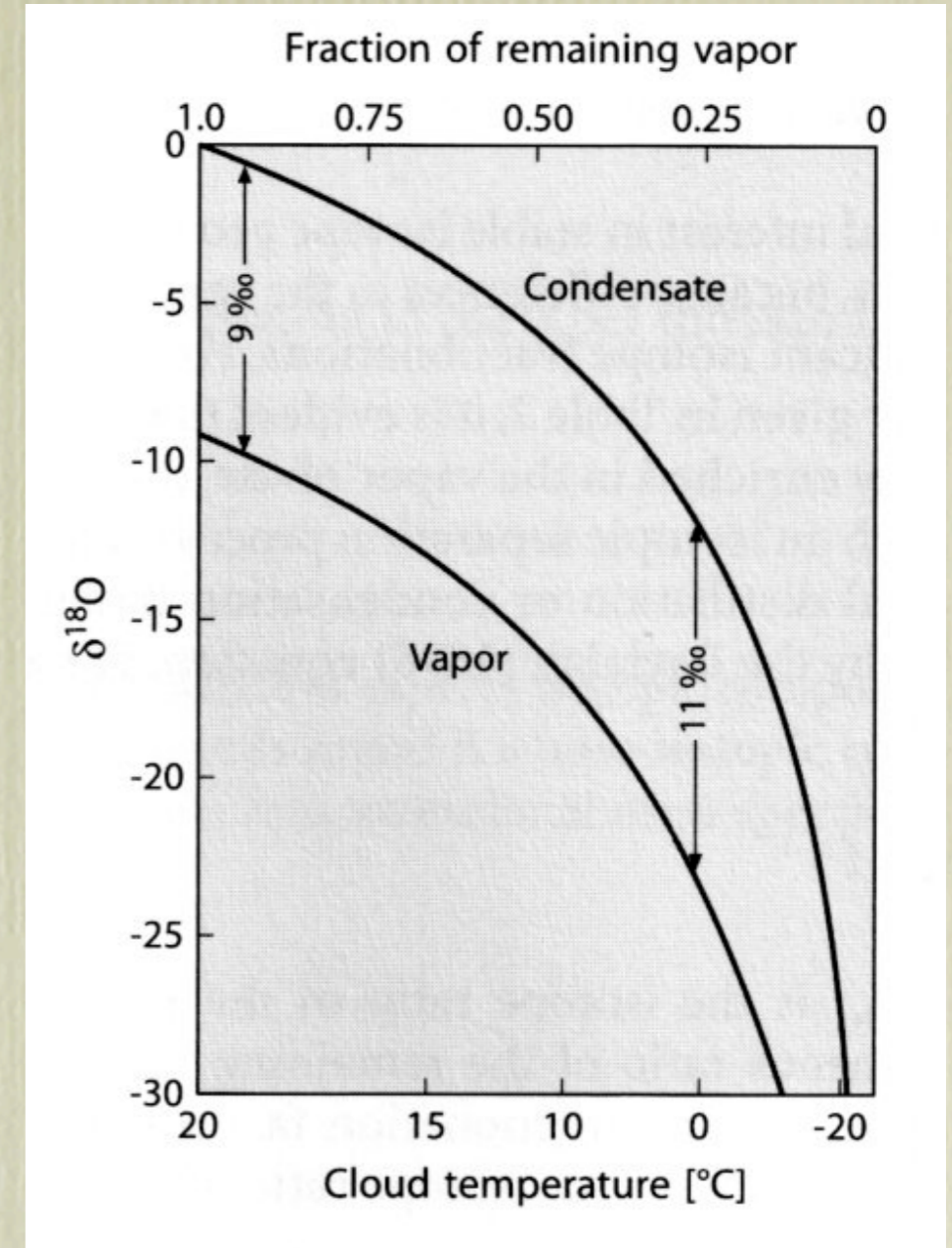
- Precipitations become lighter at higher latitudes

After Dansgaard, 1964, *Tellus*

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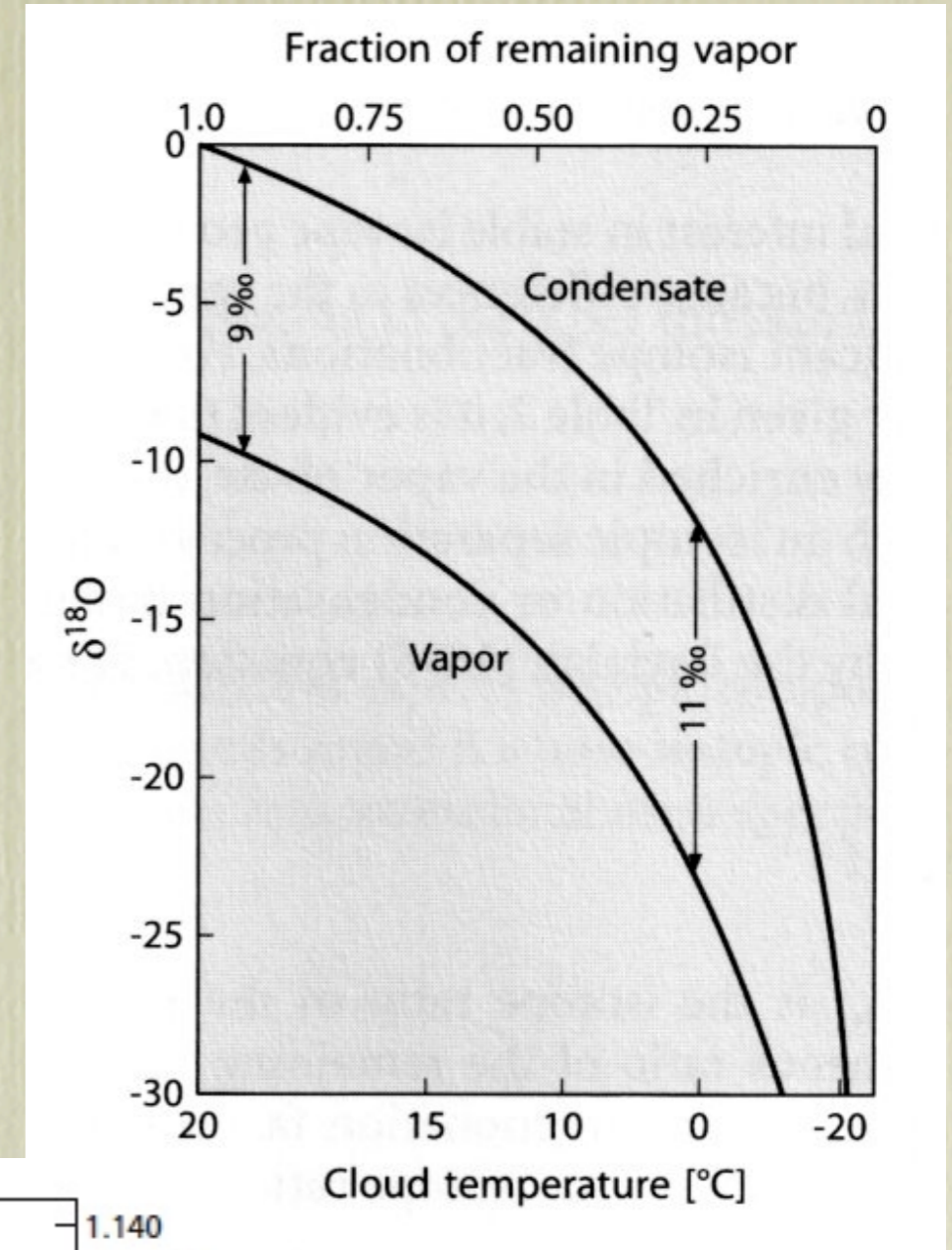
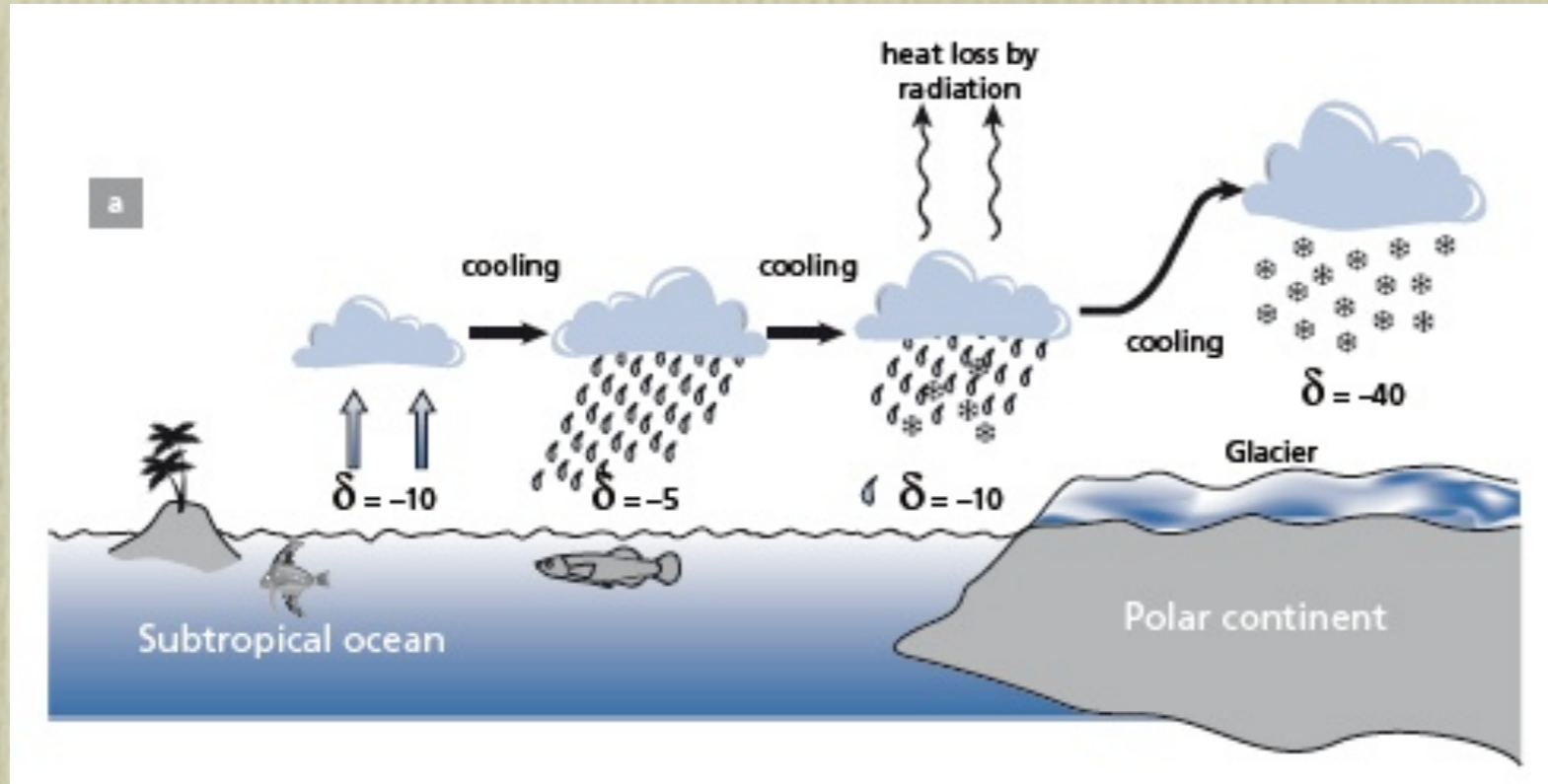


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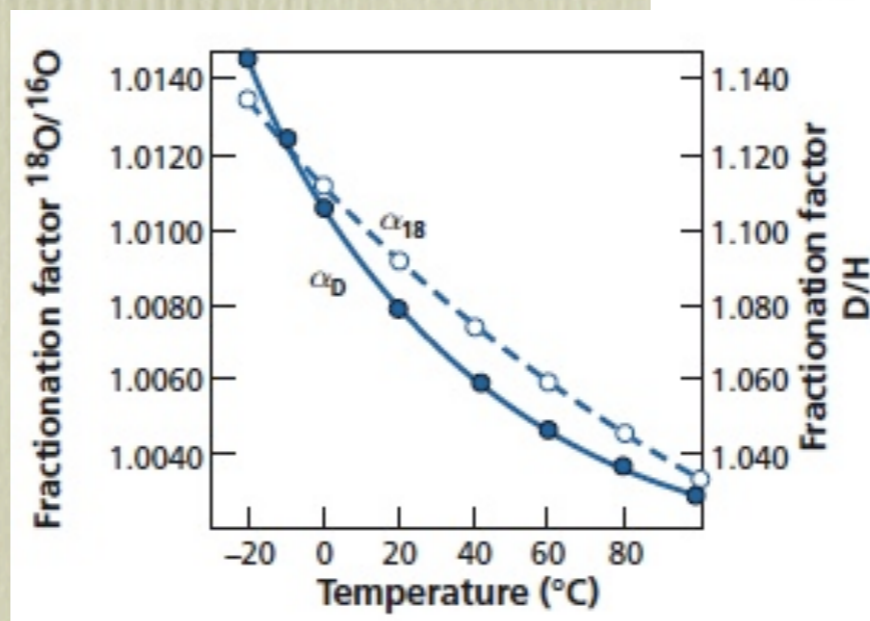


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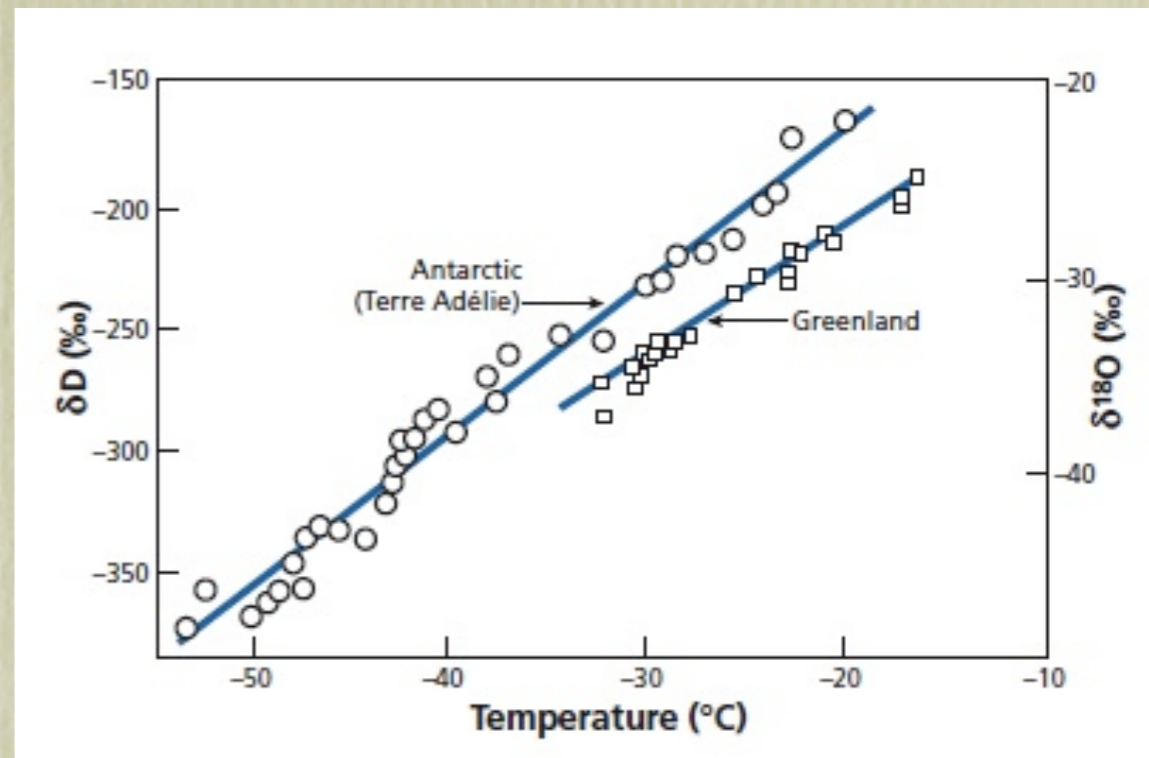


- Precipitations become lighter at higher latitudes
- α varies because there is a temperature effect
- Same story for δD



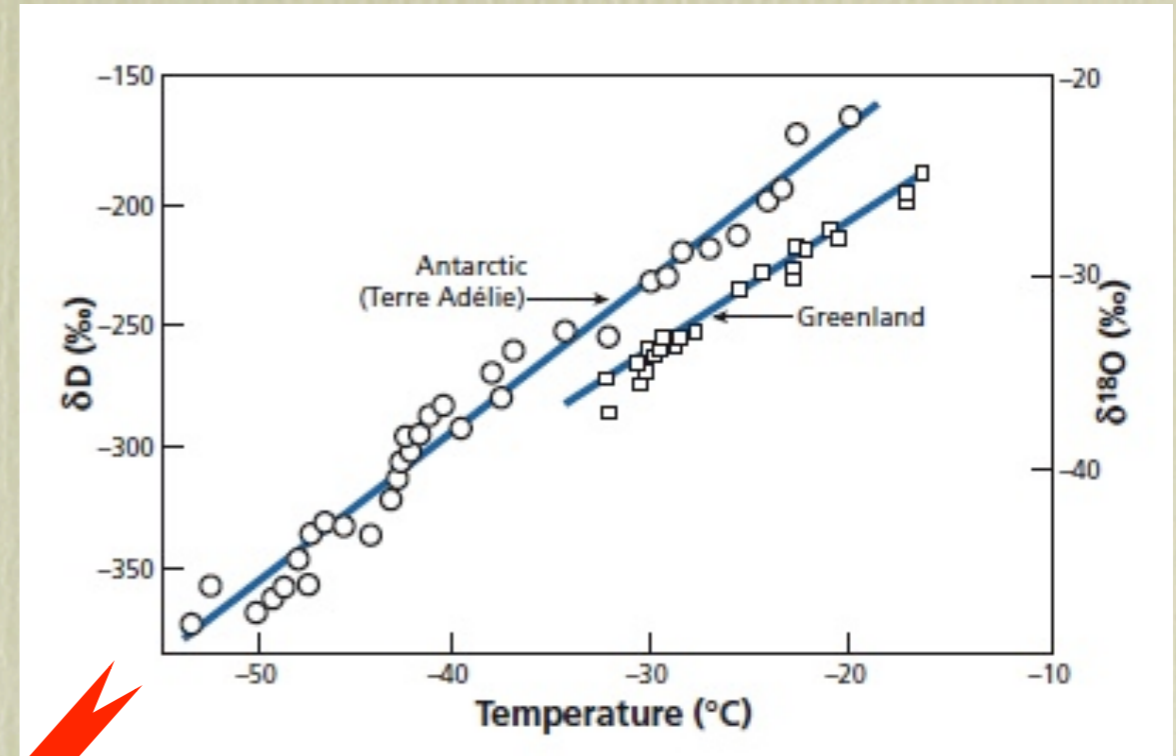
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Application for climate reconstruction using polar ice cores

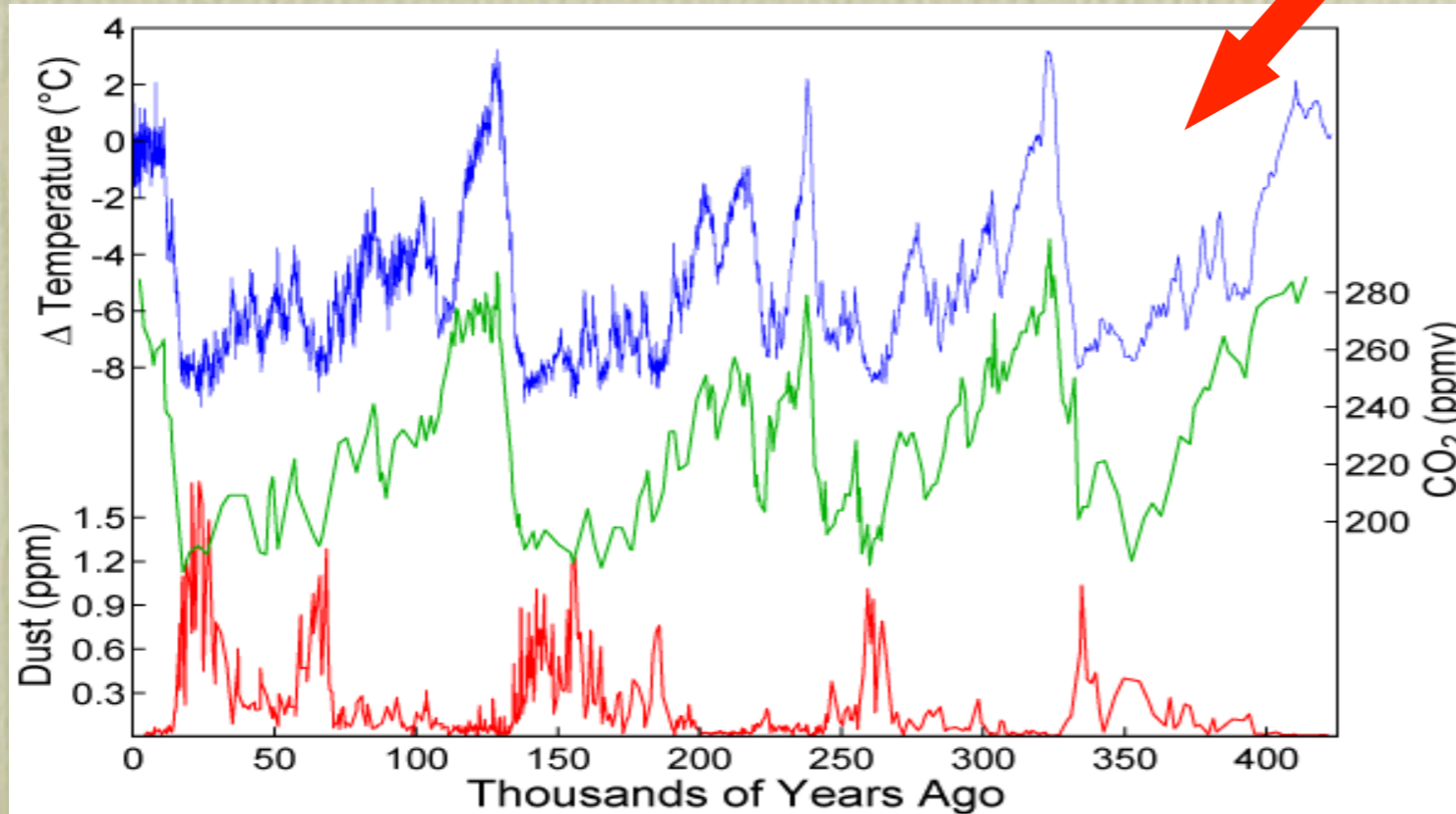


After Lorius and Merlivat, 1977;
Johnsen et al., 1989

Application for climate reconstruction using polar ice cores



After Lorius and Merlivat, 1977;
Johnsen et al., 1989



After Petit et al., 1999, Nature

- Temperature in the ice computed from δD
- Allows to show relationships between mean air temperature with other parameters, such as greenhouse gases

Suggested readings:

- Allègre, C.J., 2008. Isotope geology. Cambridge University Press, Cambridge, 512 pp.
- Hoefs, J., 2004. Stable isotope geochemistry. Springer-Verlag, Berlin, 244 pp.
- Criss, R.E., 1999. Principles of stable isotope distribution. Oxford University Press, Oxford, 254 pp.
- Reviews in Mineralogy (and Geochemistry) on Stable Isotopes (vol. 16, 1986 & vol. 43, 2001)